Reaction mechanisms in massive nuclei collisions and perspectives for synthesis of heavier superheavy elements

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Abstract We discuss a hardship in synthesis of heaviest super heavy elements in massive nuclei reactions due to the hindrance to complete fusion of reacting nuclei caused on the onset of quasifission process which strongly competes with complete fusion and due to the strong increase of fission yields along the de-excitation cascade of the compound nucleus in comparison with the evaporation residue formation. The hindrance to formation of compound nucleus and evaporation residue is determined by the characteristic of the entrance channel.

Key words Capture, Quasifission, Complete fusion, Fast fission, Fusion-fission, Evaporation residue

1 Introduction

In massive nuclei collisions, reactions evolve through various steps and different processes which determine the nature and characteristics of reaction products in dependence of the choice and conditions of reacting nuclei in the entrance channel. Therefore, different reactants which reach the same compound nucleus (CN)-characterized by the same mass number A, atomic number Z, and excitation energy E_{CN}^* do not lead to the same reaction products with the same dynamical characteristics^[1-3]. Compound nuclei formed by very different entrance channels are characterized by different angular momentum distributions at the same $\ensuremath{E_{\mathrm{CN}}^{*}}$. In this context, a very mass asymmetric reaction in the entrance channel, for which a dinuclear system (DNS) is formed after capture of reactants, mainly evolves to complete fusion (CF) and reaches the stage of CN. Then, the de-excitation behavior of CN is determined by competition between fission and particle evaporation processes. Instead, the DNS formed in a more mass

symmetric reaction in the entrance channel, undergoes strong hindrance in its evolution to CN due to the competition with the quasifission process (QF) which is the decay of DNS into two fragments. In this case, the QF yield increases and the CF rate decreases by increasing the mass and charge of nuclei in the entrance channel. Consequently, in some cases the evaporation residue (ER) cross section $\sigma_{\rm ER}$ decreases reaching values lower than 1 pb. To prove the above-mentioned statements we consider as examples of entrance channels the following reactions: 22 Ne+ 248 Cf (very mass asymmetric reaction), ²⁴Mg+²⁴⁸Cm, ³⁴S+²³⁸U (less mass asymmetric reaction), 40Ar+232Th (more mass symmetric reaction), ¹³²Sn+¹⁴⁰Ce (almost mass symmetric reaction), and ¹³⁶Xe+¹³⁶Xe (mass symmetric reaction) leading to the same ²⁷²Hs CN. In the case of the ²²Ne induced reaction the capture cross section is mainly transformed in complete fusion cross section at low energy beam^[3], and the subsequent ER formation is the dominant part of reaction products. At higher energies besides of the large contribution of QF, the FF formation competes with the ER yields at

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de-excitation of CN. In the case of the 34 S induced reaction, the QF cross section is at least one order of magnitude higher than the CN cross section because the CF is strongly hindered by the dominant role of QF, and at decay of CN the fission rate is higher than the ER process^[4]. In the case of the 40 Ar induced reaction the σ_{ER} formation is lower than 1 pb. This result is jointly caused by the dominant role of QF in comparison to fusion, and to the dominant role of the fission process in comparison to the evaporation one. Moreover, in the cases of the 132 Sn+ 140 Ce and 136 Xe+ 136 Xe reactions the ER formation completely disappears.

The competition between QF and CF processes depends on the orbital angular momentum distribution of DNS. Consequently, also the formation of the rotating and excited CN is characterized by the mass asymmetry of reactants in the entrance channel through a specific angular momentum distribution of DNS. Therefore, the same CN, which is formed by different entrance channels, characterized by the same E_{CN}^* has different angular momentum distributions. Certainly, it decays differently in competition between processes forming fission fragments (FF) or evaporation residue nuclei (ERs), along the various steps of the de-excitation cascade. Since the fission barriers (contributed by macroscopic and microscopic parts of nuclear binding energy) of CN and intermediate excited nuclei are dependent on the nuclear temperature T and angular momentum ℓ , the rates of fission fragments and ERs are sensitive to the specific dynamical properties of CN and intermediate excited nuclei determined by the used reactants in the entrance channel. Therefore, the ER cross sections decrease by increasing the angular momentum due to its influence on the rotating CN. Finally, the CN formed at the same E_{CN}^{*} by different entrance channels decays forming products (FFs and ERs) with different properties because the CN retains the dynamic peculiarities of reacting nuclei in the entrance channel. In order to give realistic estimations of the reaction product cross sections by mass symmetric or almost symmetric reactants as entrance channel, an adequate model which allows one to describe by a reliable way the complex dynamics of mechanisms

during all stages of reaction has to be developed. In fact, in the last stage of nuclear reaction, the formed CN may de-excite by fission (producing fusion-fission fragments) or by emission of light particles. The reaction products that survive fission are ERs. The registration of ER is a clear evidence of the CN formation, but in case of reactions with massive nuclei, generally, the knowledge of the ERs formation only is not enough to determine the complete fusion cross section and to understand the dynamics of the de-excitation cascade of CN if the true fission fragments are not correctly taken into account. On the other hand, the correct identification of an evaporation residue nucleus by the observation of its decay chain does not assure if the target material contains other isotopes of the nucleus under consideration. In fact, for example, in the case of the ⁴⁸Ca+²⁴⁹Cf reaction, the identification of the ²⁹⁴118 nucleus as the evaporation residue of the ²⁹⁷118 CN after the emission of 3 neutrons (see the experiment reported in Ref.[5]) cannot assure that the collected events corresponding to the ²⁹⁴118 nucleus are obtained only due to the mentioned reaction leading to the formation of the ²⁹⁷118 CN. The interaction of the ⁴⁸Ca projectile with the ²⁵⁰Cf isotope in target should be considered in order to take into account the contribution of the ⁴⁸Ca+²⁵⁰Cf reaction to the ²⁹⁴118 ER formation because the target material inevitably contains the ²⁵⁰Cf isotope too. In fact, in this last case, the ⁴⁸Ca+²⁵⁰Cf reaction (forming the ²⁹⁸118 CN) leads to the same ²⁹⁴118 evaporation residue nucleus after emission of 4 neutrons from CN. This effect depends on the E_{CN}^* excitation energy of CN which is determined by the collision energy $E_{\rm c.m.}$. In addition, the use of some assumptions about the reaction mechanisms leading to the formation of the observed fission-like fragments, does not allow for sure correct determination of the fusion-fission contribution in the case of overlapping of the mass fragment distributions due to different processes (quasifission, fast fission fusion-fission) $^{[1,3,6,7]}$. The exigency importance to have a multiparameter and sensitive model is strongly connected with the requirement to reach reliable results and with the possibility to give

(1)

reliable estimations of perspectives for the synthesis of superheavy elements (SHE).

Model and formalism 2

By using the DNS model^[8], the first stage of reaction is the capture formation of a DNS after full momentum transfer of the relative motion of colliding nuclei into a rotating and excited nuclear system. In the deep inelastic collisions DNS is formed but the full momentum transfer does not occur. Therefore, the deep inelastic collisions are not capture reactions.

The partial capture cross section at a given energy $E_{\rm cm}$ and orbital angular momentum ℓ is determined by Eq.(1).

$$\sigma_{\rm cap}^\ell(E_{\rm c.m.}) = \pi \lambda^2 P_{\rm cap}^\ell(E_{\rm c.m.}) \qquad (1)$$
 where the capture probability $P_{\rm cap}^\ell(E_{\rm c.m.})$ —equal to 1 or 0 for a given $E_{\rm c.m.}$ energy and orbital angular momentum ℓ in dependence on the result of dynamical calculations—is the path of collision trapped into the nucleus-nucleus potential well or not, respectively, after dissipation of part of the initial relative kinetic energy and orbital angular momentum^[1,9]. Our calculations showed that, depending on the center-of-mass system energy $E_{\rm c.m.}$, there can be "window" in the orbital angular momentum for capture with respect to the conditions described in Refs.[1,9]. The quasifission process competes with formation of CN. This process occurs when the DNS prefers to break down into two fragments instead of transforming into the fully

$$\sigma_{\text{fus}}(E_{\text{c.m.}}; \beta_p, \alpha_T) = \sum_{\ell=0}^{\ell_f} (2\ell + 1)$$

$$\times \sigma_{\text{cap}}(E_{\text{c.m.}}; \ell, \beta_p, \alpha_T) P_{\text{CN}}(E_{\text{c.m.}}; \ell, \beta_p, \alpha_T).$$
(2)

equilibrated CN. The fusion excitation function is

determined by product of the partial capture cross section $\sigma_{\text{cap}}^{\ell}$ and the fusion probability P_{CN} of DNS, at

various $E_{\rm cm}$ values:

Obviously, the quasifission cross section is defined by

$$\sigma_{\text{qfis}}(E_{\text{c.m.}}; \beta_p, \alpha_T) = \sum_{\ell=0}^{\ell_f} (2\ell+1)$$

$$\times \sigma_{\text{can}}(E_{\text{c.m.}}; \ell, \beta_p, \alpha_T) (1 - P_{\text{CN}}(E_{\text{c.m.}}; \ell, \beta_p, \alpha_T)).$$
(3)

For more specific details and descriptions of the model see Refs.[1,7,9-11]. In order to show the sensitivity of our model, we present in Fig.1 the calculated $P_{\rm CN}$ fusion probability as a function of the orbital angular

momentum ℓ , at excitation energies E_{CN}^* =49 MeV (dashed line) and 63 MeV (full line) of the ²⁰²Pb CN in the ⁴⁸Ca+¹⁵⁴Sm reaction.

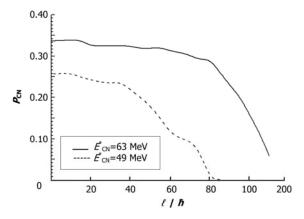


Fig.1 The $P_{\rm CN}$ fusion probability calculation vs. the orbital angular momentum ℓ for the $^{48}{\rm Ca}{}^{+154}{\rm Sm}$ reaction, at two different $E_{\rm CN}^*$ values of the $^{202}{\rm Pb}$ CN.

Figure 1 shows how much the P_{CN} fusion probability changes with the ℓ value at a fixed E_{CN}^* excitation energy of CN, and how much the $P_{\rm CN}$ trend of changes at two different $E_{\rm CN}^*$ values. Therefore, the methods that do not take into account in calculation the dependence of the $P_{\rm CN}$ fusion probability on the collision energy $E_{\rm c.m.}$, angular momentum ℓ , and on the orientation angles of the axial symmetry axes of deformed reacting nuclei cannot reach reliable values of the fusion cross section.

The fast fission cross section is calculated by summing the contributions of the partial waves corresponding to the range $\ell_f \le \ell \le \ell_d$ leading to the formation of a mononucleus where the fission barrier B_f is zero^[7] in such range of ℓ , and therefore the system promptly decays into two fragments:

$$\sigma_{\text{fastfis}}(E_{\text{c.m.}}; \beta_p, \alpha_T) = \sum_{\ell=\ell_f}^{\ell_d} (2\ell+1) \times \sigma_{\text{cap}}(E_{\text{c.m.}}; \ell, \beta_p, \alpha_T) (P_{\text{CN}}(E_{\text{c.m.}}; \ell, \beta_p, \alpha_T)).$$
(4)

The capture cross section is equal to the sum of the quasifission, fusion, and fast fission cross sections^[6]:

$$\sigma_{\text{cap}}^{\ell}(E_{\text{c.m.}};\beta_{p},\alpha_{T}) = \sigma_{\text{qfiss}}^{\ell}(E_{\text{c.m.}};\beta_{p},\alpha_{T}) + \sigma_{\text{fis}}^{\ell}(E_{\text{c.m.}};\beta_{n},\alpha_{T}) + \sigma_{\text{fistfis}}^{\ell}(E_{\text{c.m.}};\beta_{n},\alpha_{T}).$$

$$(5)$$

It is clear that the fusion cross section includes the cross sections of evaporation residues and fusion-fission products^[7]. The ER cross section is calculated by the advanced statistical code^[12-14] that takes into account the damping of the shell correction in the fission barrier as a function of nuclear temperature and orbital angular momentum in determination of the survival probability

$$\sigma_{\text{ER}(x)}(E_x^*) = \sum_{\ell=0}^{\ell} (2\ell+1)\sigma_{(x-1)}^{\ell}(E_x^*)W_{\text{sur}(x-1)}(E_x^*,\ell).$$
 (6)

We are able to calculate mass- and angle-distribution of quasifission and fusion-fission fragments, anisotropy of the fission fragment angular distribution and the dependence of cross sections, Coulomb barrier, intrinsic fusion barrier and quasifission barrier as a function of the orientation angle of the symmetry axes of colliding nuclei (see Refs.[3,7,15]). In Fig.2 we present, as an example, the mass distribution of quasifission fragments for the ⁴⁸Ca+¹⁵⁴Sm reaction.

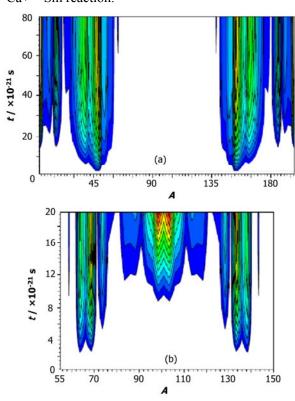


Fig.2 (Color online) (a) Mass distributions of the quasifission products yield in the 48 Ca+ 154 Sm reaction at $E_{\rm c.m.}$ =140 MeV as a function of the lifetime of the dinuclear system formed at capture stage. (b) Mass distributions of the quasifission product yields in the 48 Ca+ 154 Sm reaction at $E_{\rm c.m.}$ =160 MeV as a function of the lifetime of the dinuclear system.

In many cases, in dependence on the entrance channel peculiarities, the mass distributions of the fusion-fission, quasifission, and fast fission fragments can overlap^[3,10]. As a result, the real difficulties arise

in the analysis of experimental data in order to identify the true yields of fragments according to corresponding processes in heavy-ion collisions. Fig.2 shows that at lower $E_{\rm c.m.}$ energy the mass distribution of quasifission products populates the asymmetric mass region at any lifetime value of DNS (Fig.2a), while at higher $E_{\rm c.m.}$ energy it is also populated the symmetric mass region for longer DNS lifetimes (Fig.2b). The lifetime, in fact, of an excited DNS for a given reaction depends on the initial collision energy $E_{\rm c.m.}$ and angular momentum distribution values. Therefore, the DNS during its evolution can evolve to complete fusion (fusion process) or can decay into two fragments (quasifission process).

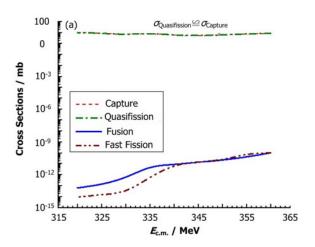
The competition between these two processes is related to the values of intrinsic fusion barrier $B_{\rm fus}^*$ and quasifission barrier $B_{\rm qf}^{[1,2,16]}$ depending on the peculiarities of reacting nuclei, beam energy and angular momentum distribution.

3 Comparison between the $^{136}\mathrm{Xe+}^{136}\mathrm{Xe}$ and $^{24}\mathrm{Mg+}^{248}\mathrm{Cm}$ reactions leading to $^{272}\mathrm{Hs}$ CN

In order to check if any projectile and target combination can always lead to the complete fusion of reactants (having an enough high energy beam to overcome the Coulomb barrier) and synthesis of the wanted SHE, we consider the case of the ¹³⁶Xe+¹³⁶Xe mass symmetric reaction which would lead to the ²⁷²Hs CN. By using the procedure presented in the previous Section, for this reaction, the results are shown in Fig.3. Fig.3a shows the capture, quasifission, fusion and fast fission cross sections vs. $E_{c.m.}$ energy. And Fig.3b shows the fusion probability $P_{\rm CN}$ in the same explored $E_{c.m.}$ energy range. As one can see the capture cross section for the ¹³⁶Xe+¹³⁶Xe reaction is about 10 mb in the explored energy range while the fusion cross section leading to the ²⁷²Hs CN ranges between 10^{-4} and 10^{-1} pb (with a fusion probability of about 10^{-14} – 10^{-11}) in the same $E_{\rm cm}$ interval. By the present investigation we can conclude that the evaporation residue cross section is much lower than 10⁻¹⁰ pb. Such a value practically means that no synthesis event of reacting nuclei occurs.

For a comparison with the results for the last reaction, as shown in Fig.4(a), the results obtained in

this work for the mass asymmetric ²⁴Mg+²⁴⁸Cm reaction leading to the same ²⁷²Hs CN, where in lower-medium $E_{c.m.}$ energy range, the fusion process is dominated in the reaction dynamics. At high energy, the quasifission process prevails. Fig.4(b) shows the ER cross sections. From the comparison of the results presented in Figs.3 and 4, we can conclude that the fusion-fission cross section in the ²⁴Mg+²⁴⁸Cm reaction at E_{CN}^* of about 55 MeV ($E_{c.m.}$ =192 MeV) is 150 mb, while the results for above-mentioned fusion-fission cross section with the mass symmetric distribution in the 136Xe+136Xe reaction is lower than some pb (because the reaction dynamics is completely dominated by the quasifission For this symmetric reaction, fusion-fission yield is at the same E_{CN}^{\star} excitation energy of 55 MeV ($E_{c.m.}$ =355 MeV), at least 10^{-12} times lower than the one obtained by the ²⁴Mg+²⁴⁸Cm reaction.



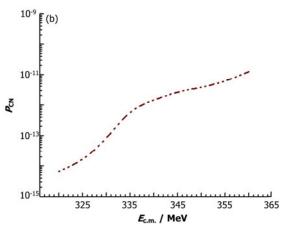
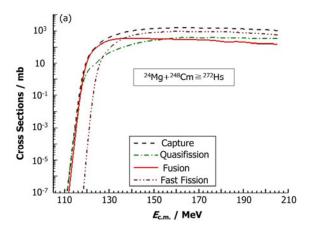


Fig.3 (Color online) Capture, quasifission, fusion and fast fission cross sections (panel (a)), and the $P_{\rm CN}$ fusion probability (panel (b)), vs. the $E_{\rm c.m.}$ energy of the $^{36}{\rm Xe} + ^{136}{\rm Xe}$ reaction.



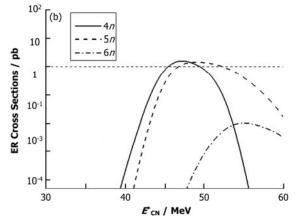


Fig.4 (Color online) Capture, quasifission, fusion and fast fission cross sections (panel (a)) vs. the energy, and the individual evaporation residue cross sections (panel (b)) versus the $E_{\rm CN}^*$ excitation energy of CN, for the $^{24}{\rm Mg}{+}^{248}{\rm Cm}$ reaction. In panel (b) the results of our calculation by using the masses and barriers of Refs.[17,18] are reported.

4 Study on superheavy nuclei and perspectives for heavier superheavy elements

In order to estimate the realistic possibilities of synthesis of SHE by massive nuclei reactions, we performed calculations of the ER cross sections for set of reactions forming fissile compound nuclei with $Z\geq 100$ at the same excitation energy ($E_{CN}^* \sim 37$ MeV).

In Table 1, we present the set of elements by various entrance channels with different charge (mass) asymmetry parameters. It is interesting to observe and analyze the overall trend of the fusion probability $P_{\rm CN}$ and the evaporation residue yields for various reactions as a function of the charge $Z_{\rm CN}$ and of the parameter $z=(Z_1\times Z_2)/(A_1^{1/3}+A_2^{1/3})$ in order to draw some useful indications on the possible reactions leading to heavy nuclei with $Z_{\rm CN} \ge 100$ and particularly on reactions leading to SHE with $Z_{\rm CN} \ge 120$. Fig.5 shows the fusion probability $P_{\rm CN}$ for the reactions

listed in Table 1 as a function of the charge Z_{CN} at excitation energy $E_{CN}^* \sim 37$ MeV. As shown in Fig.5, $P_{\rm CN}$ slowly decreases with $Z_{\rm CN}$ but strongly decreases for more symmetric reactions in entrance channel leading to the same Z_{CN} . The trend of P_{CN} for the same investigated reactions appears more clear if we report the calculated $P_{\rm CN}$ as a function of the parameter $z=(Z_1\times Z_2)/(A_1^{1/3}+A_2^{1/3})$ representing the Coulomb barrier of interacting nuclei in the entrance channel if we divide this z parameter to the r_0 nuclear parameter that is able to calculate the radius of each nucleus $(R=r_0A^{1/3})$ (see Fig.7 from Ref.[19]). In this last case, the values of $P_{\rm CN}$ reported at the given values of $Z_{\rm CN}$ (108, 118, 120, 122, 122, 124 and 126) represent different fusion probabilities for various entrance channels of reactions leading to the same Z_{CN} . The fusion probability P_{CN} strongly decreases by increasing the z parameter and by decreasing the charge (mass) asymmetry parameter of reactions in the entrance channel. The hindrance to fusion increases for more symmetric reactions and for higher Coulomb barriers of reactions in entrance channel. The evaporation residues after neutron emission only from the de-excitation cascade of CN can be observed for reactions with z parameter lower than the value of about 200. For reactions with values of z parameter included in the range about 200-235 the observation of residues is at limit (or it appears to be a very problematic task) of the current experimental possibilities. For reactions with z higher than 235 it is impossible to observe ER of CN after neutron emission only. We report in Table 2 that the results obtained for the investigated reactions leading to CN with Z=120, 122, 124 and 126, at $E_{CN}^* \sim 37$ MeV.

Table 1 Listed reactions are reported as a function of the charge Z_{CN} of CN (if it can be reached), and the parameter $z=(Z_1\times Z_2)/(A_1^{1/3}+A_2^{1/3})$ related to the Coulomb barrier of reacting nuclei in the entrance channel

Reactions	$Z_{\rm CN}$	Z	Reactions	$Z_{\rm CN}$	Z
$^{16}O + ^{238}U$	100	84	⁸⁶ Kr+ ²⁰⁸ Pb	118	286
⁴⁸ Ca+ ²⁰⁸ Pb	102	172	$^{132}Sn + ^{174}Yb$	120	328
⁵⁰ Ti+ ²⁰⁸ Pb	104	188	64Ni+238U	120	253
136 Xe $+^{136}$ Xe	108	284	⁵⁸ Fe+ ²⁴⁴ Pu	120	242
⁵⁸ Fe+ ²⁰⁸ Pb	108	218	⁵⁴ Cr+ ²⁴⁸ Cm	120	229
⁴⁸ Ca+ ²²⁶ Ra	108	181	¹³² Sn+ ¹⁷⁶ Hf	122	337
26 Mg+ 248 Cm	108	125	⁵⁴ Cr+ ²⁴⁹ Cf	122	234
$^{48}\text{Ca} + ^{243}\text{Am}$	115	193	$^{132}Sn+^{186}W$	124	343
⁴⁸ Ca+ ²⁴⁸ Cm	116	194	⁵⁸ Fe+ ²⁴⁹ Cf	124	251
⁴⁸ Ca+ ²⁴⁸ Bk	117	196	84Kr+232Th	126	307
⁴⁸ Ca+ ²⁴⁹ Cf	118	198	⁶⁴ Ni+ ²⁴⁹ Cf	126	267

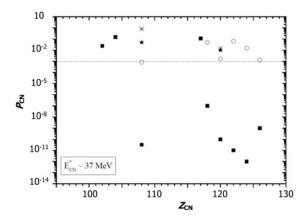


Fig.5 Fusion probability $P_{\rm CN}$ calculated at the same excitation energy $E_{\rm CN}^*$ 37 MeV versus charge $Z_{\rm CN}$ for the reactions listed in Table 1. The different symbols (full squares, open circles, full stars and asterisks) are related to our calculated $P_{\rm CN}$ values in respect of the reactions listed in Table 1. The $P_{\rm CN}$ values are higher for higher mass asymmetric reactions.

Table 2 Reactions leading to compound nuclei with $Z_{\rm CN}$ =120, 122, 124 and 126. As a function of the parameter z related to the Coulomb barrier in the entrance channel. $\sigma_{\rm ER}$ is the ER cross section after the neutron emission only from the de-excitation cascade of CN. $P_{\rm res/cap}$ is the ratio between the yields of evaporation residue $\sigma_{\rm ER}$ and the capture $\sigma_{\rm cap}$

Reactions	$Z_{\rm CN}$	Z	$\sigma_{ m ER/mb}$	$P_{ m res/cap}$
⁵⁴ Cr+ ²⁴⁸ Cm	120	229	1.05×10^{-10}	0.30×10^{-11}
⁵⁸ Fe+ ²⁴⁴ Pu	120	242	5.40×10^{-12}	1.70×10^{-14}
$^{64}Ni+^{238}U$	120	253	3.10×10^{-15}	1.40×10^{-16}
⁵⁴ Cr+ ²⁴⁹ Cf	122	234	1.40×10^{-10}	1.30×10^{-12}
⁵⁸ Fe+ ²⁴⁹ Cf	124	251	1.61×10^{-15}	1.80×10^{-17}
64Ni+249Cf	126	267	4.40×10^{-20}	6.50×10^{-22}

We estimated that only for the SHE with $Z_{\rm CN}$ =120 it is possible to observe evaporation residues by reactions with z parameter lower than 230. The possibility of obtaining the heaviest 302119 and 305120 SHEs by using the ⁴⁸Ca beam in the ⁴⁸Ca+²⁵⁴Es and ⁴⁸Ca+²⁵⁷Fm reactions, respectively, is restricted by difficulties in obtaining enough thick targets of ²⁵⁴Es and ²⁵⁷Fm because the other Es and Fm isotopes are radioactive with shorter lifetimes. Therefore, in order to reach heavier SHE, reactions with beams heavier than ⁴⁸Ca (as for example ⁵⁰Ti, ⁵⁴Cr, ⁵⁸Fe, ⁶⁴Ni and other heavier projectiles) against the above-mentioned actinide targets should be used. But, unfortunately, the evaporation residues cross sections strongly decrease by decreasing the charge (mass) asymmetry of reactants in the entrance channel.

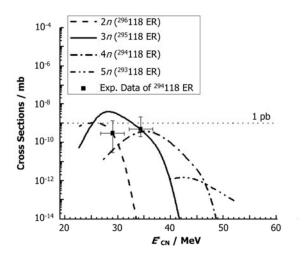


Fig.6 Individual evaporation residue excitation functions after emission of 2 (dashed line), 3 (full line), 4 (dash-dotted line) and 5 (dash-double dotted line) neutrons from the ²⁹⁷118 CN in the reaction of ⁴⁸Ca+²⁵⁰Cf. The experimental data (full squares) of the ²⁹⁴118 ER formation cross section obtained from Ref.[5].

The first experiments which were performed at Flerov Laboratory of Nuclear Reaction of Joint Institute for Nuclear Reaction (58 Fe+ 244 Pu[20]) and at GSI of Darmstadt (64 Ni+ 238 U and 54 Cr+ 248 Cm[21], and 50 Ti+ 249 Cf[22]) to explore the synthesis of the Z=120 SHE did not identify any event of the expected SHE. In our previous papers (Refs.[10,15]), we presented results of calculation about the above-mentioned reactions which could lead to the Z=120 SHE, but we found values of the evaporation residue cross sections lower than 0.1 pb. Predictions of other authors are approximately near this value[$^{23-27}$]. Therefore, it is necessary to improve the experimental conditions in order to be able to reach measurements of cross sections of the order of fb.

Moreover, we also studied four reactions induced by ⁴⁸Ca on the ²⁴⁹⁻²⁵²Cf targets in order to analyze the effect of mass number and structure properties of nuclei in the entrance channel on the capture, quasifission, and complete fusion processes. The study and comparison of capture cross sections allows us to reveal the sensitivity of the model and results on the dynamical effects of the entrance channel (for results see Fig.1 of Ref.[28]), while the determination and analysis of the evaporation residue cross sections for the four reactions reveal the influence of the different structure of the formed ²⁹⁷⁻³⁰⁰118 superheavy compound nuclei in the ⁴⁸Ca+²⁴⁹⁻²⁵²Cf reactions with different neutron rich

targets. In the following Figs.6 and 7, for example, the ER excitation functions obtained for the ⁴⁸Ca+^{249,250}Cf reactions, respectively, by using the masses and barriers of Refs.[17,18]. We also investigated the formation of the heaviest evaporation residue nuclei from the ^{299,300}118 CNs which are formed in reactions induced by collision of the ⁴⁸Ca projectiles with the heaviest accessible ^{251,252}Cf actinide targets, and the results are comparable with the ones obtained for the reactions on the ^{249,250}Cf targets.

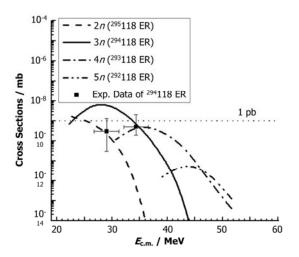


Fig.7 As Fig.6, but for the ⁴⁸Ca+²⁴⁹Cf reaction.

By analyzing the 2, 3, 4 and 5 neutron emission channels along the de-excitation cascade of compound nuclei formed in the 48Ca+249,250Cf reactions, we studied the possibilities of synthesizing the ²⁹²⁻²⁹⁶118 ER nuclei. In addition, by considering the experimental conditions nowadays available in Laboratories, the more convenient and accessible reaction channels of observing evaporation residue nuclei are the 3 and 4 neutron emission channels in the ⁴⁸Ca+²⁴⁹⁻²⁵²Cf reactions at beam energies corresponding to the $E_{\rm CN}^*$ =25-40 MeV interval. By comparing the results of our analysis regarding the study of the ⁴⁸Ca+^{249,250}Cf reactions with the data obtained in the experiment of Ref.[5] regarding the observation of the ²⁹⁴118 evaporation residue nucleus, we conclude that the better description of the experimental results is that the observed ²⁹⁴118 synthesis events^[5], registered at two different beam energies, are contributed by the 3n-channel in the ⁴⁸Ca+²⁴⁹Cf reaction and 4*n*-channel in the ⁴⁸Ca+²⁵⁰Cf reaction, due to the inevitable presence of the 250Cf isotope in the ²⁴⁹Cf enriched target. Moreover, the

comparison of results obtained for the ER nuclei in the 48 Ca+ 252 Cf studied reaction suggest the use of one target only constituted of all the Cf isotopes having longer lifetimes. This is more convenient either for the procedure of its preparation or in analysis of one experiment only. In fact, it is possible to observe and study a wide set of ER nuclei formed by 2n, 3n, 4n, and 5n emission channels, only changing the 48 Ca beam energy E_{lab} in the range of 235–260 MeV.

5 Conclusion

At present time, it is a problematic task to measure ER cross sections of SHE with Z=120, and this is also impossible for reactions with z parameter higher than about 240. Then, mass symmetric reactions with z>240, as for example $^{136}Xe+^{136}Xe$ reaction (where z=284), cannot form ER nuclei because that reaction does not give sufficient fusion cross section. It is impossible to obtain SHE's with Z>120 by complete fusion reactions since the z parameter is higher than 240. In reaction induced by ⁴⁸Ca beam it is impossible to obtain ER nuclei higher than ²⁹⁸118 by using Cf targets. Instead, by using a mixture of Cf isotopes as target, it is possible to explore by one experiment only the ²⁹⁴⁻²⁹⁸118 ER cross sections in the 25-40 MeV excitation energy range. From the study of the present investigated systematics on reactions for the superheavy formations, we understand the role of the mass symmetry parameter of entrance channel on the fusion probability of reaction and evaporation residue yields. Regarding the results of the investigated reactions leading to the formation of compound nuclei with $Z_{\text{CN}}=120$, 122, 124, and 126 we affirm that it is still possible to reach and observe ER nuclei of the Z=120 SHE by reactions with z parameter of about 230, while it is a very doubtful venture to synthesize the Z=122 SHE by reactions with z parameter of about 234, or higher, by the current experimental resources and methods of observing evaporation residues. It appears out of every possibility to observe evaporation residue of SHE by reactions with z parameter in the entrance channel higher than 240. Therefore, it is impossible to form the Z=124 and Z=126 superheavy reactions. nuclei above-mentioned by the Consequently, it is an unrealizable dream to think of performing the $^{132}Sn+^{208}Pb$ (with z=373) and

 132 Sn+ 249 Cf (with z=431) reactions in order to reach the 340 132 and 381 148 SHE, respectively, and by mass symmetric reactions like 139,149 La+ 139,149 La (with z=317 and 306, respectively) in order to synthesize heavy and superheavy elements because of the absolute dominant contribution of the quasifission process after capture, and the fast fission process presents at stage of the already small probable formation of complete fusion.

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